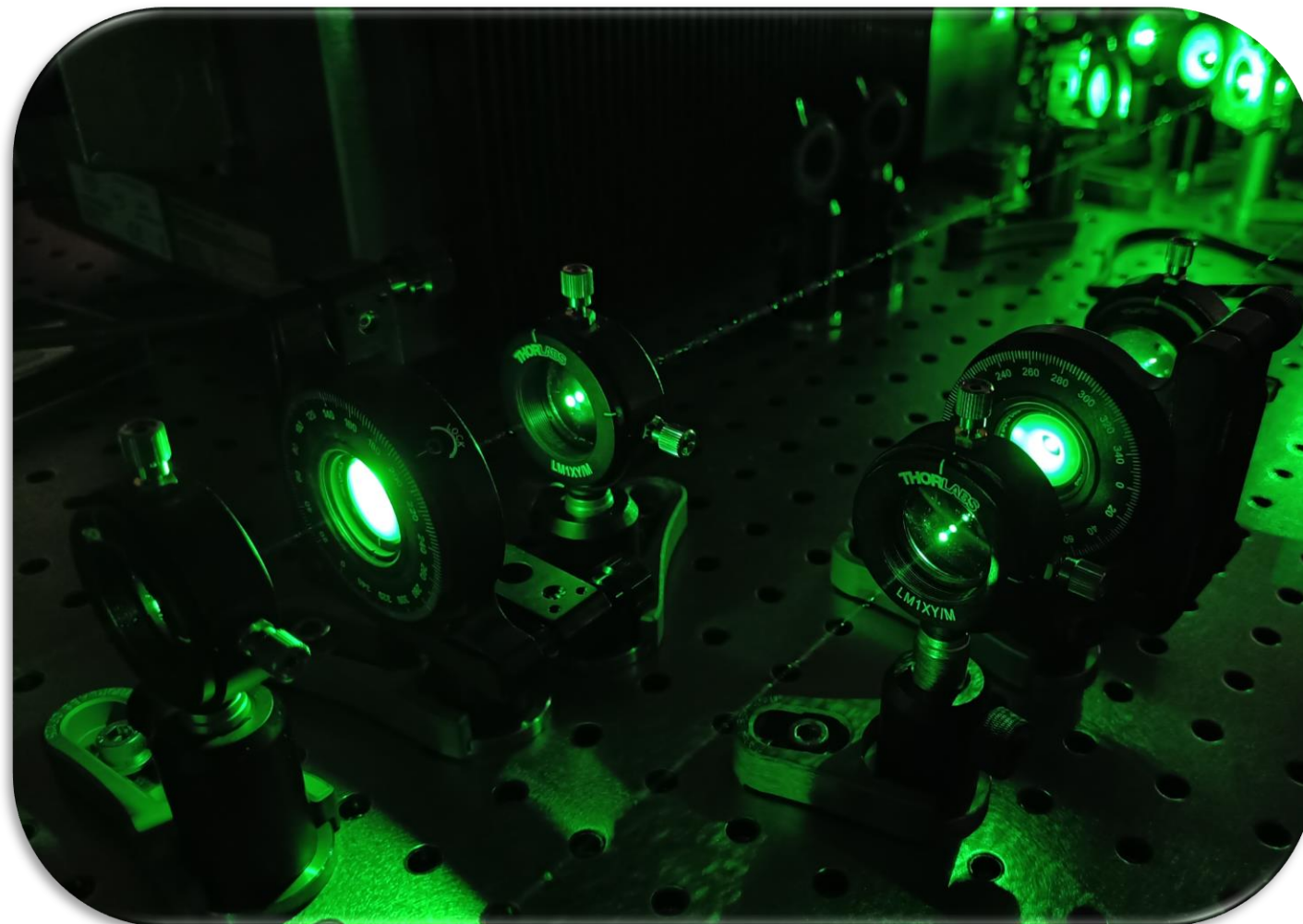


# (UV) Dual Comb Spectroscopy: Ultra-High Resolution Molecular Spectra in Milliseconds

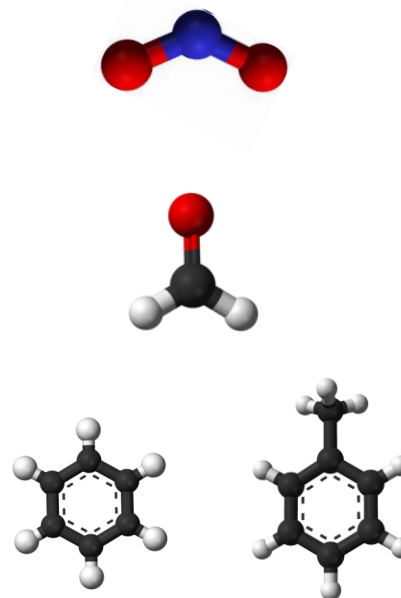
Elias Ehl<sup>1,2</sup>, Birgitta Bernhardt<sup>1</sup>,  
Luca Fossati<sup>2</sup>

<sup>1</sup>Graz University of Technology,  
<sup>2</sup>OeAW Space Research Institute

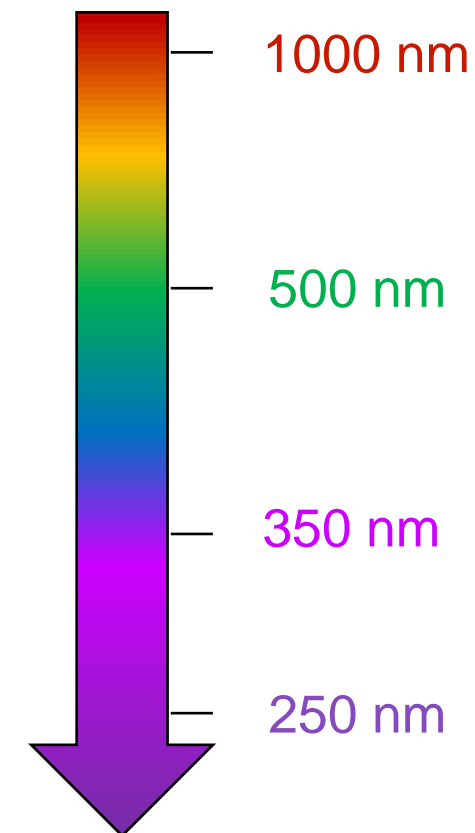


# Outline

- Dual Comb Spectroscopy
- Nitrogen Dioxide (VIS)
- Formaldehyde (NUV)
- Benzene & Toluene (DUV)



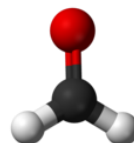
Wavelength Region



# Why High Resolution & High Speed?

Current methods:

or



**Slow & accurate**

Resolve rotational structure



Time-dependent measurements



**Fast & low resolution**

Resolve rotational structure



Time-dependent measurements



**Dual Comb Spectrometer**

Resolve rotational structure



Time-dependent measurements

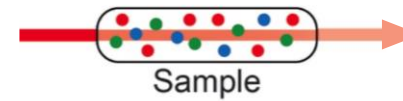
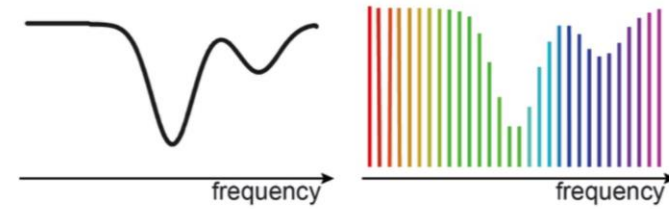
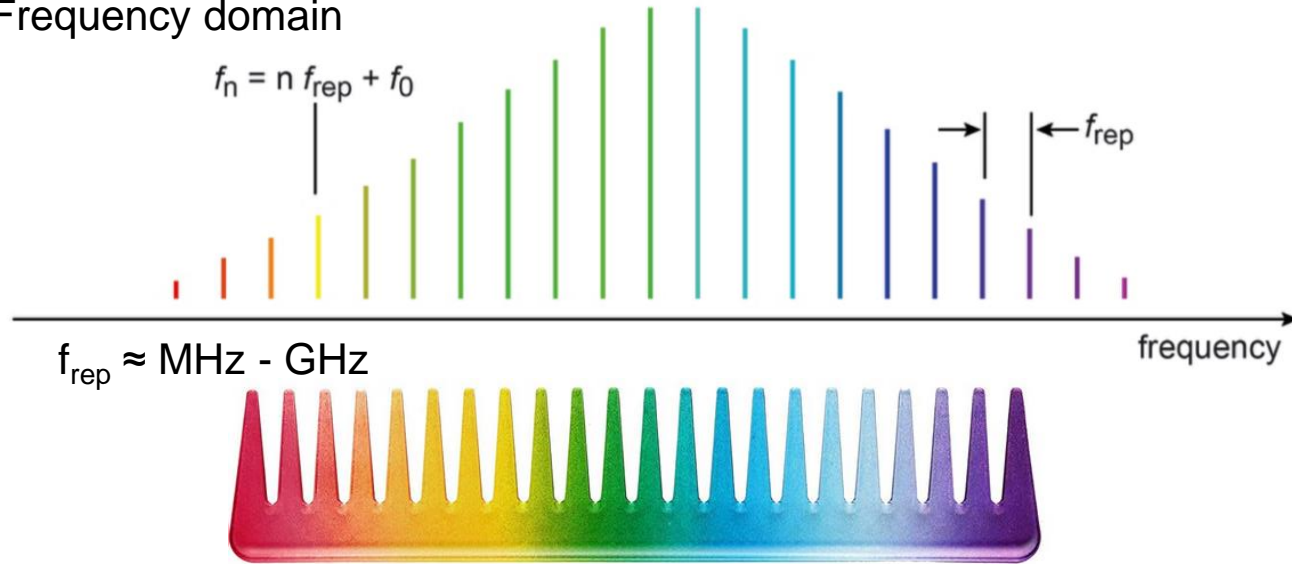




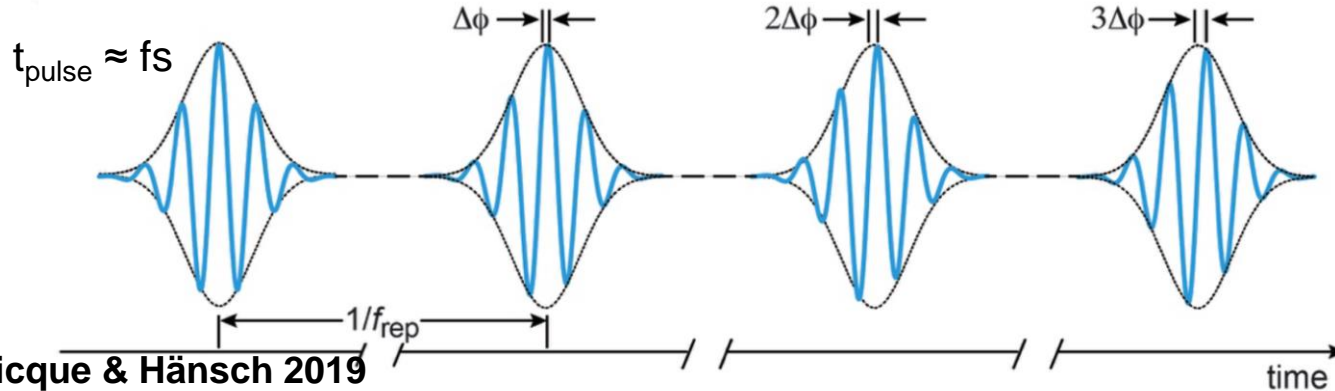
Glauber, Hall & Hänsch, 2005

# Frequency Combs

Frequency domain



Time domain

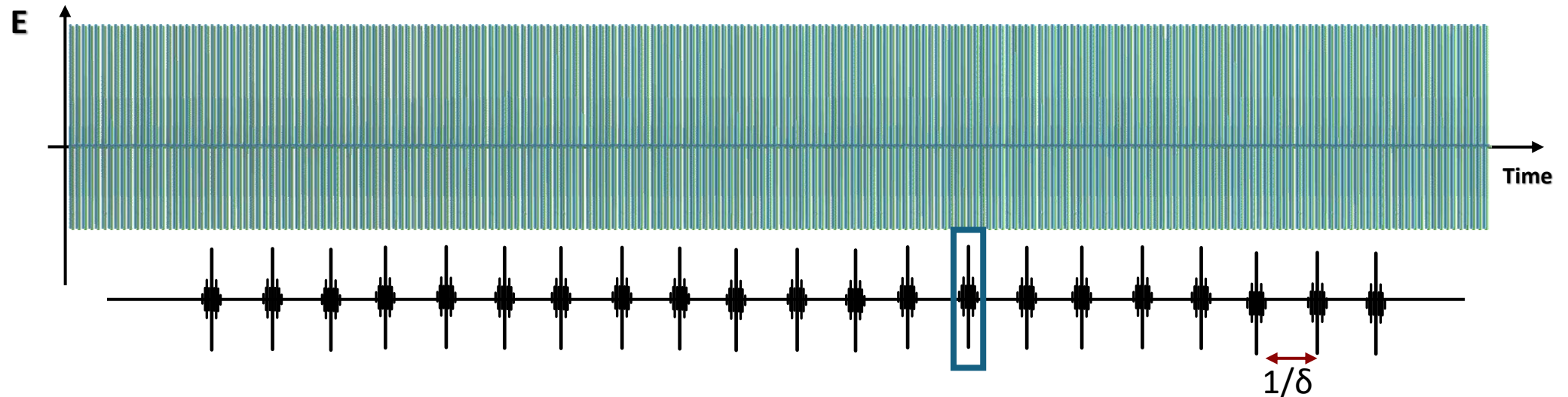
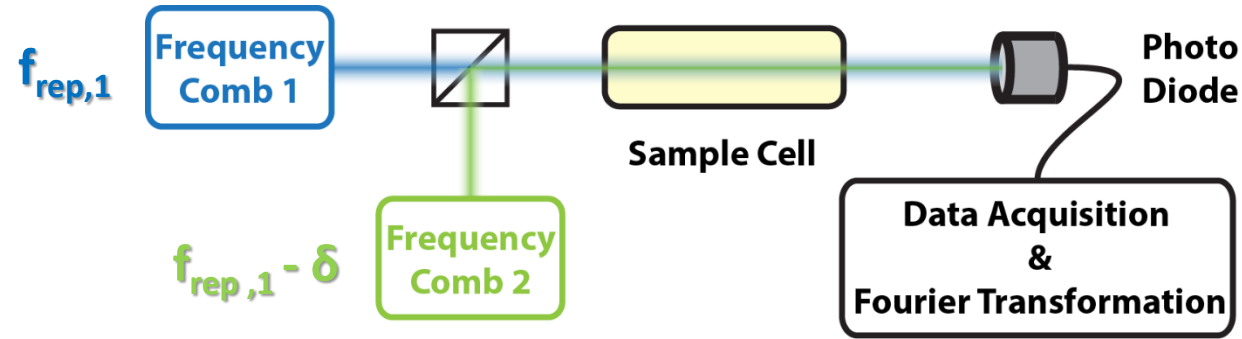


High absolute precision!  $10^{-14}$



Image credit: ESO

# Dual Comb Spectroscopy (DCS)



**Frequency down-conversion**  
 $\sim 10^{14} \text{ Hz} \rightarrow \sim 10^6 \text{ Hz}$

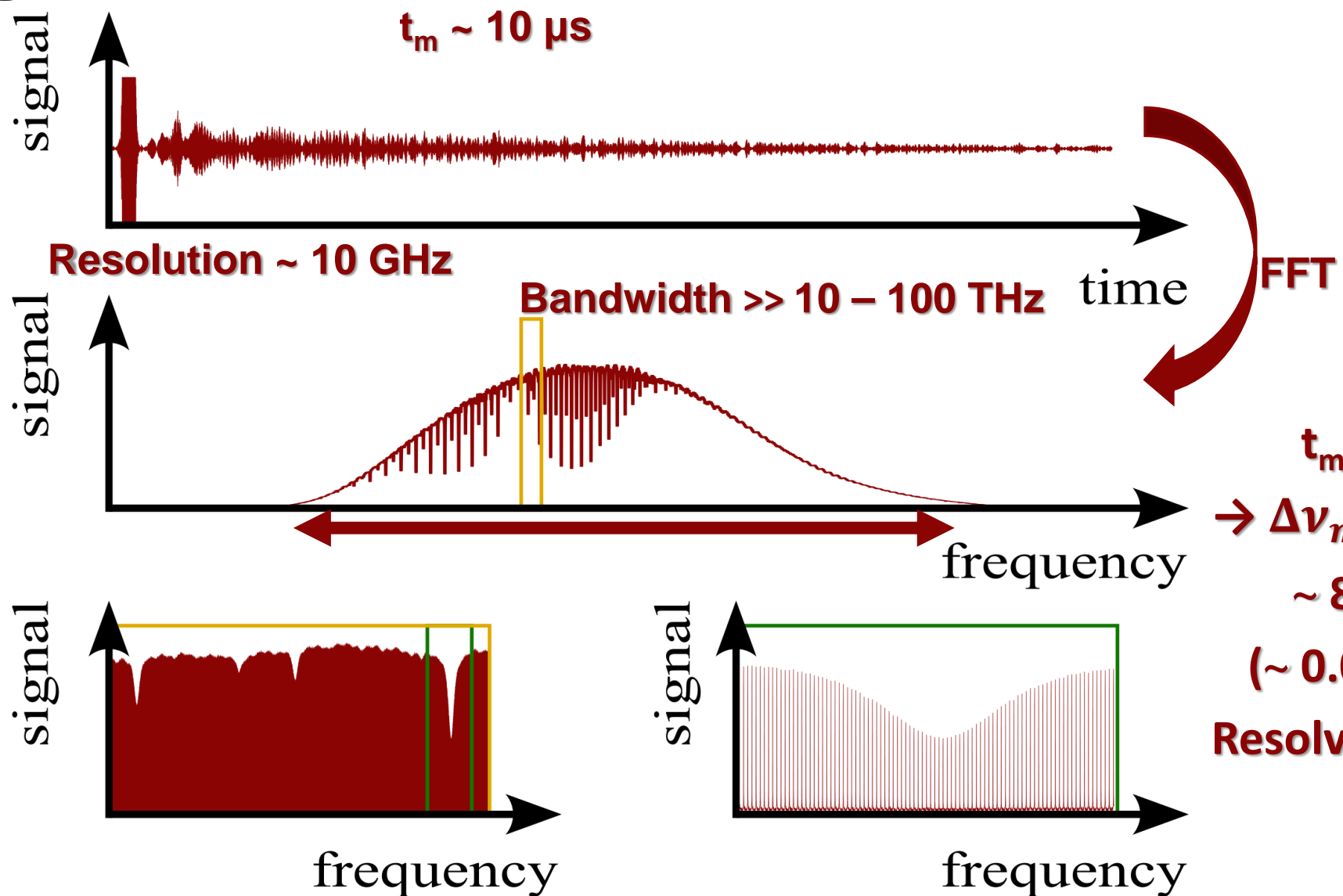
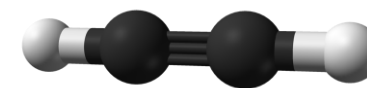


**Static spectrometer**



**Ultrashort acquisition times**

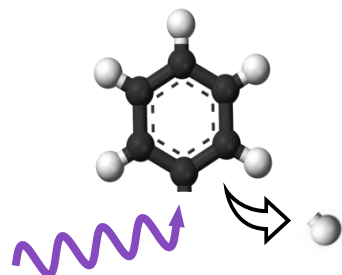
# Strengths of DCS



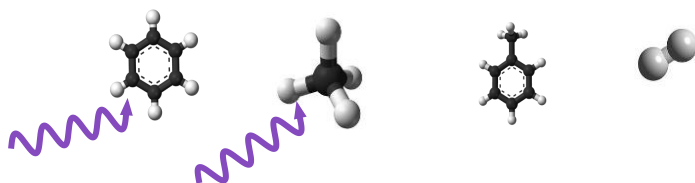
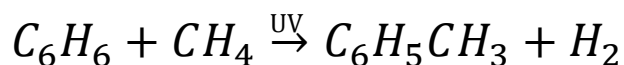
# Why UV & Visible?

## Atmospheric chemistry

- Photodissociation

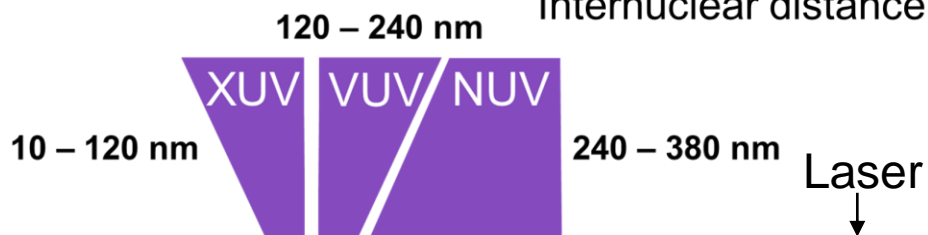
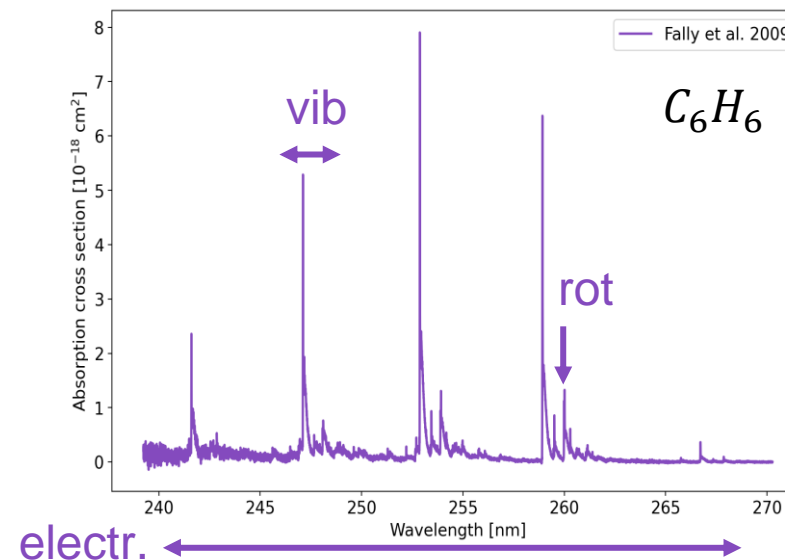
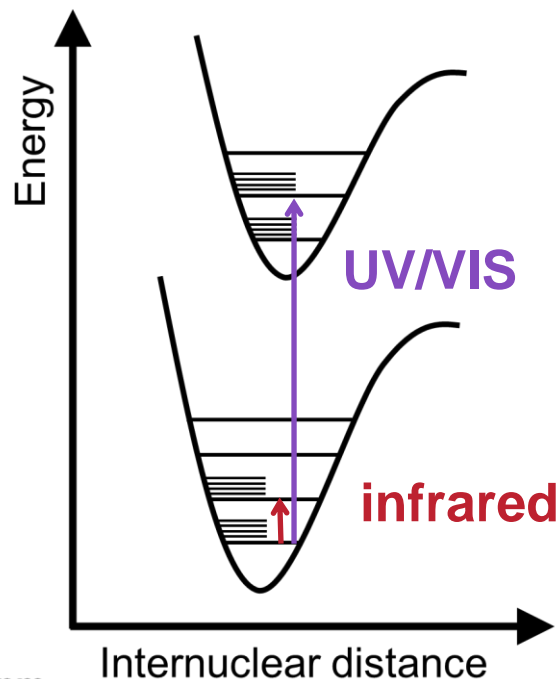


- Forming new molecules

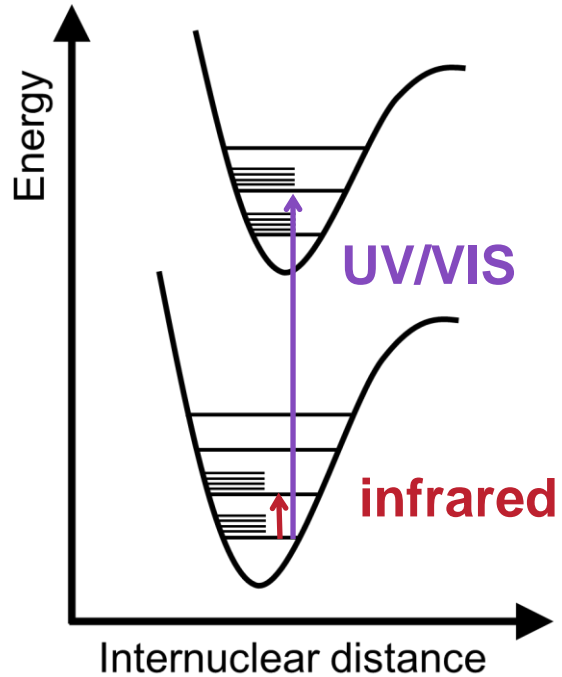


## Uniquely identify molecules

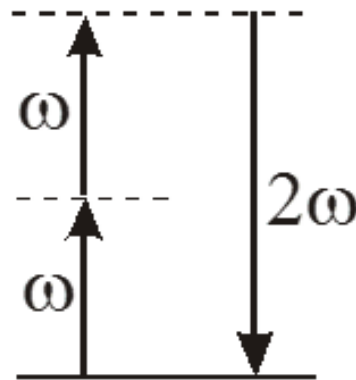
- Vibrational, rotational and electronic transitions



# Accessing New Wavelengths



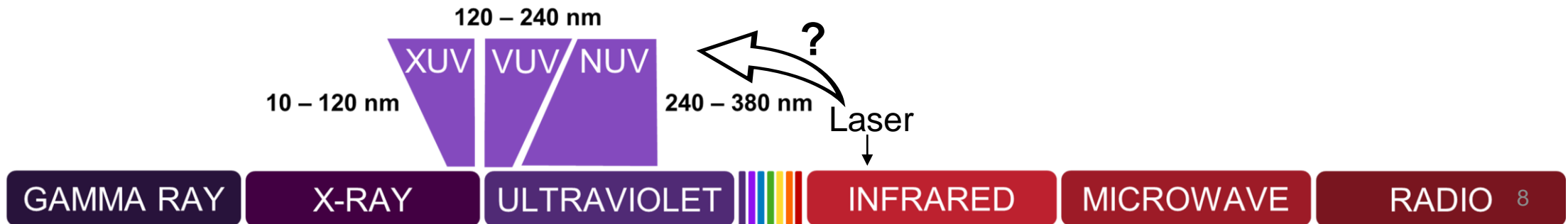
- Nonlinear frequency upconversion



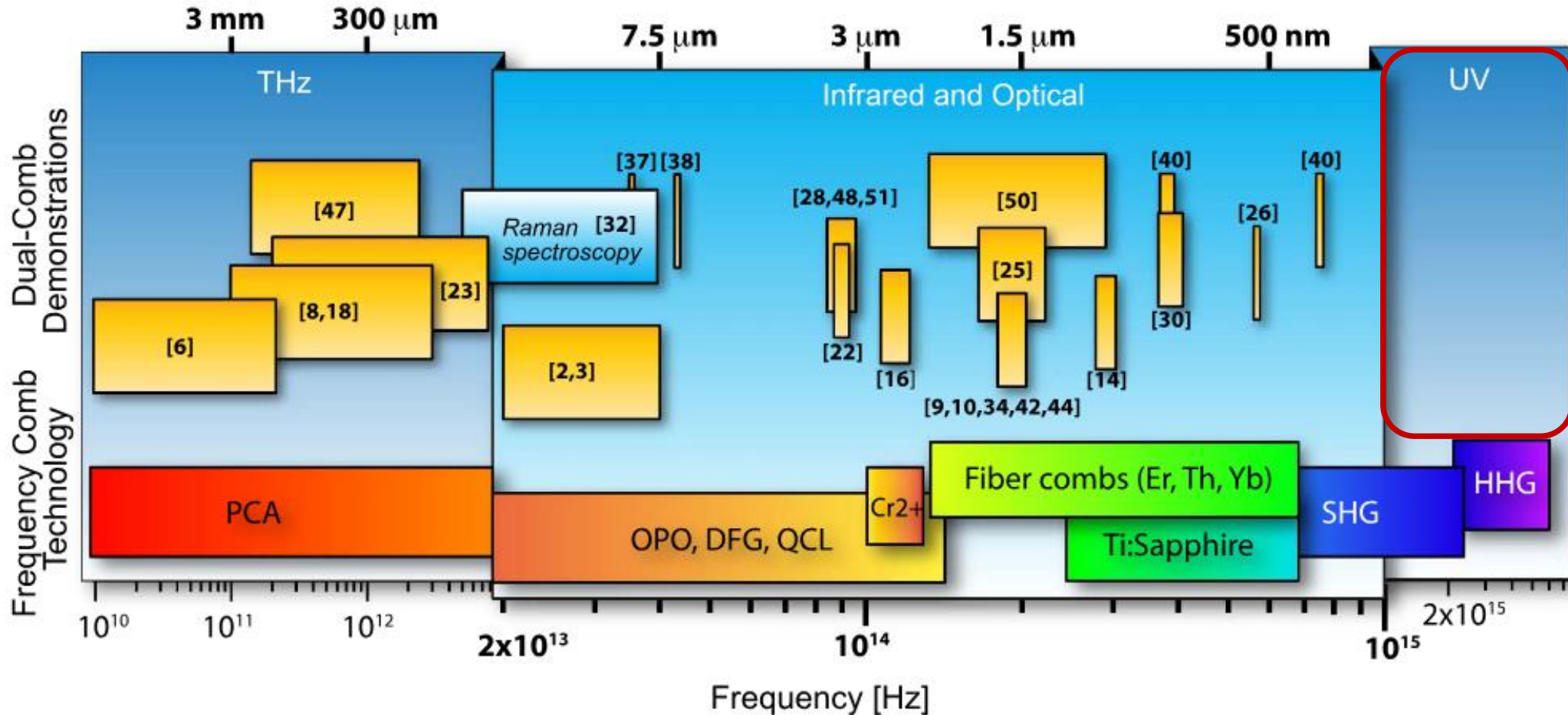
1050 nm → 525 nm **2<sup>nd</sup> Harmonic (SHG)**

1050 nm → 350 nm **3<sup>rd</sup> Harmonic (THG)**

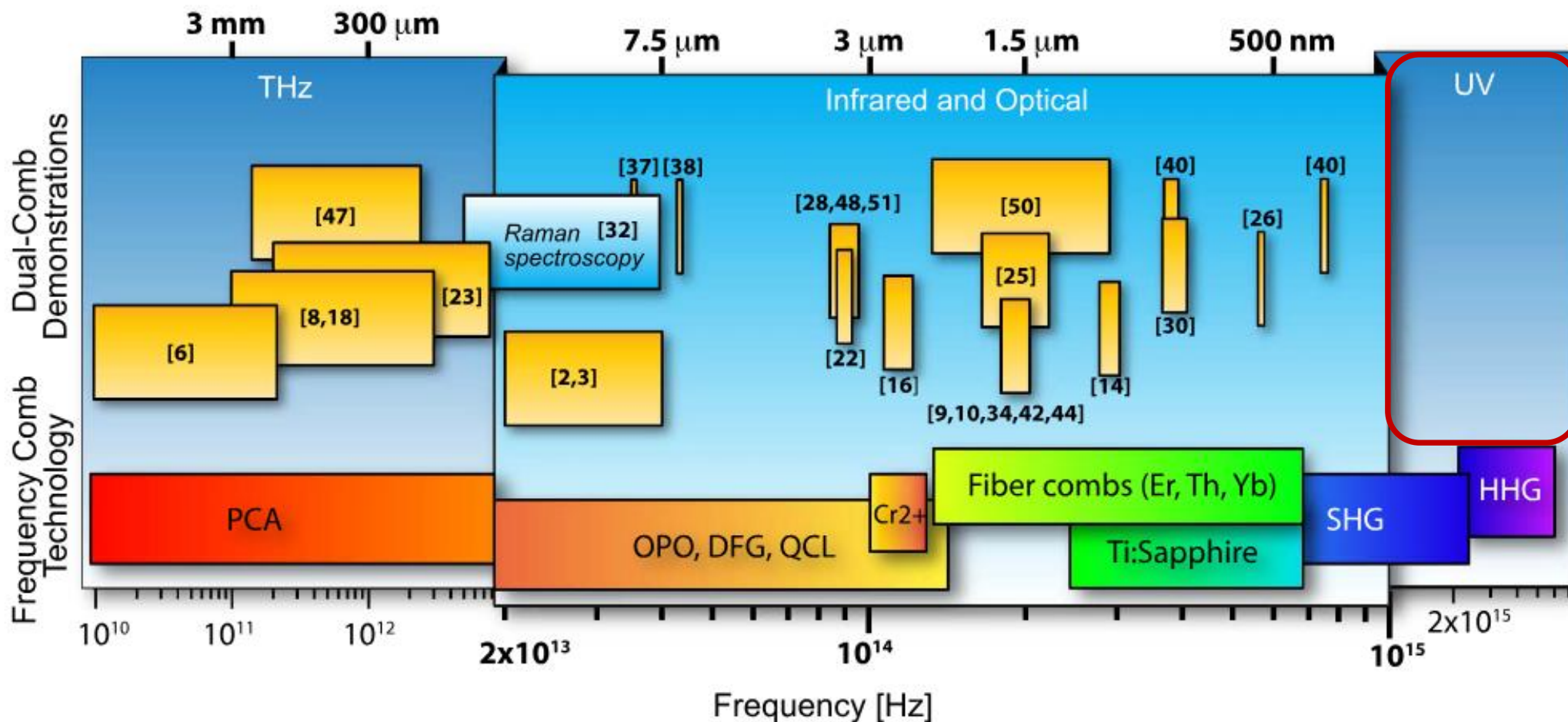
1050 nm → 262 nm **4<sup>th</sup> Harmonic (FHG)**



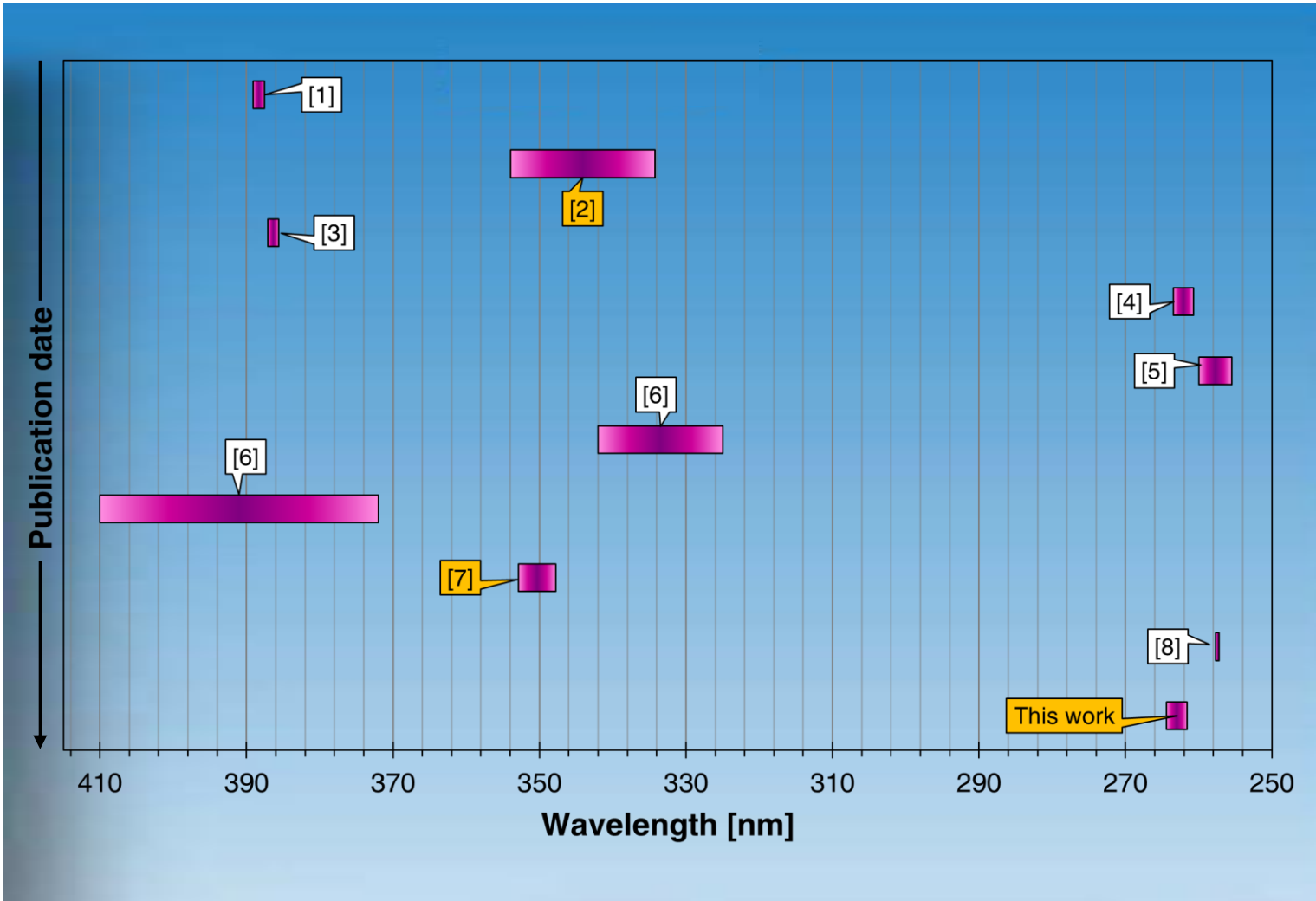
# DCS State-of-the-Art



# UV DCS State-of-the-Art

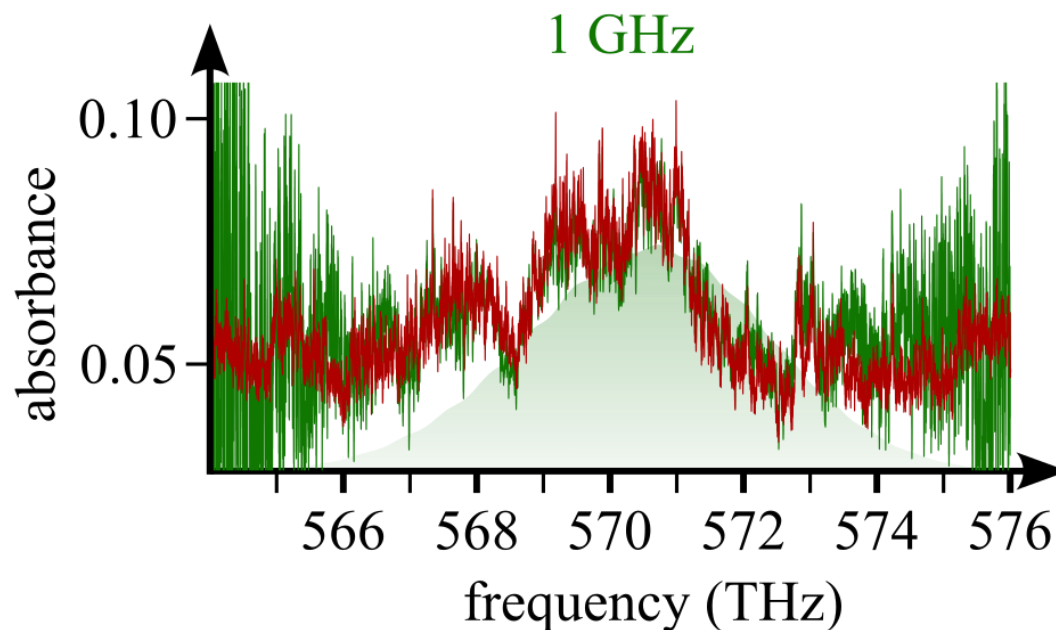
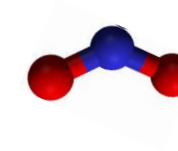


# UV DCS State-of-the-Art



[1]	Xu et al. 2024 DOI: 10.1038/s41586-024-07094-9	MPQ Garching, GER
[2]	Fuerst et al. 2024 DOI: 10.1364/OPTICA.516783	TU Graz, AUT
[3]	Chang et al. 2024 DOI: 10.1364/OL.515776	NIST, USA
[4]	McCauley et al. 2024 DOI: 10.1364/OPTICA.516851	University of Arizona, USA
[5]	Li et al. 2025 DOI: 10.34133/ultrafastscience.0087	Aerospace Information Research Institute, CHN
[6]	Muraviev et al. 2024 DOI: 10.1364/OPTICA.536971	University of Central Florida, USA
[7]	Fuerst et al. <b>subm.</b> DOI: 10.48550/arXiv.2501.07350	TU Graz, AUT
[8]	Hofer et al. 2025 DOI: 10.1088/1361-6455/add4d7	HSU Hamburg, GER
This work	Ehl et al. <b>in prep</b>	TU Graz, AUT

# VIS (526 nm) – NO<sub>2</sub>



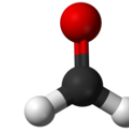
Alexander Eber

$$\lambda [nm] = \frac{3 \cdot 10^5}{f [THz]}$$

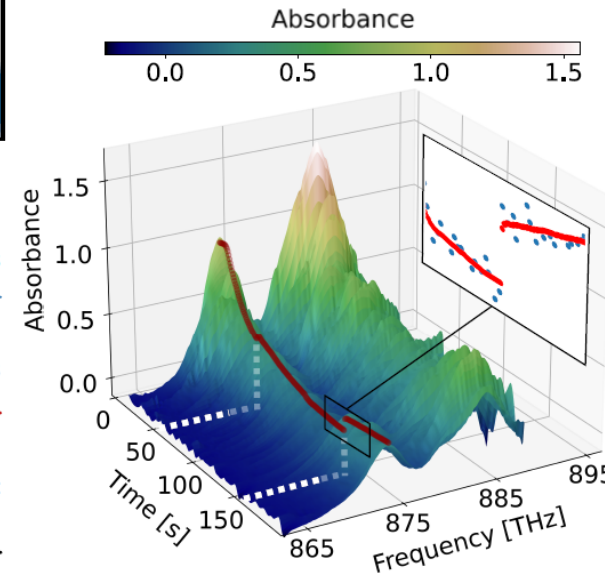
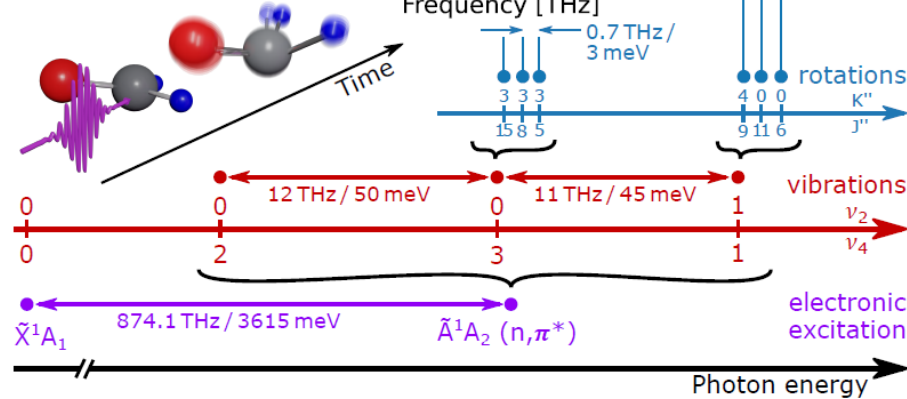
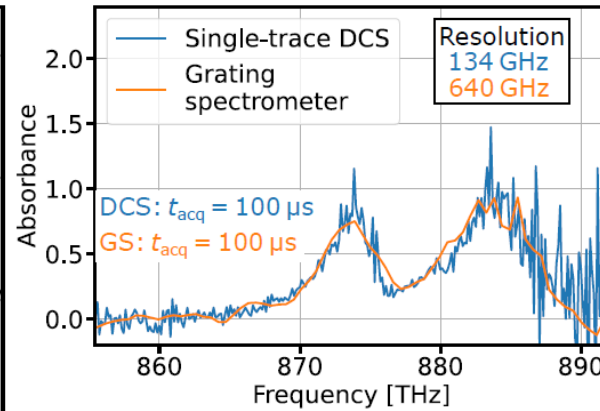
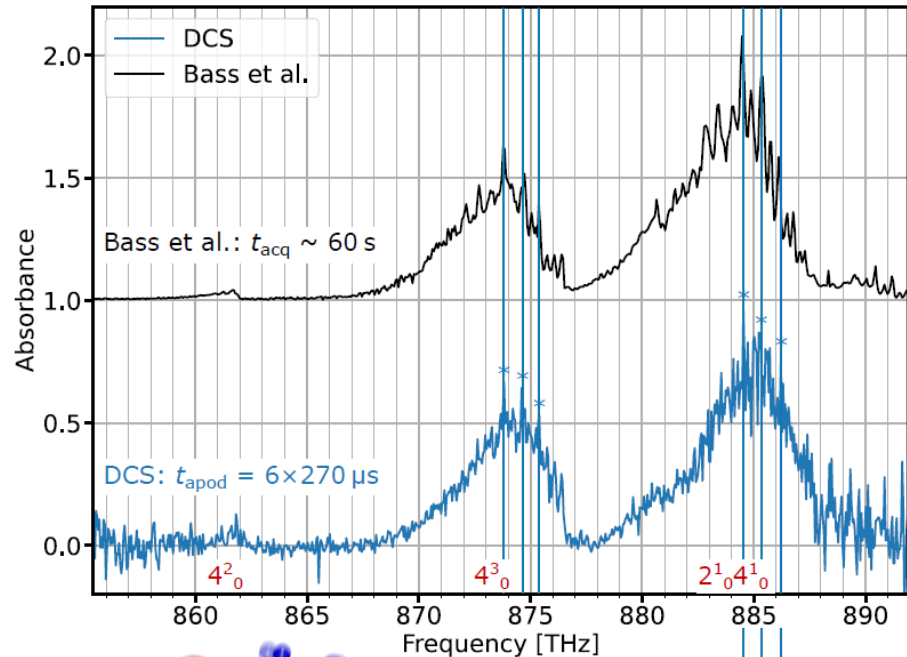
Doppler Width: 1 GHz

— DCS, 1 GHz — Nizkorodov et al.

	Resolution	Acquisition Time
DCS (1 GHz)	1000 MHz	1 s
FTS (Nizkorodov et al. 2004)	890 MHz	80 min



# NUV (350 nm) – CH<sub>2</sub>O



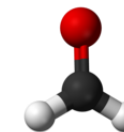
**Comparable resolution  
in a fraction of the time**



Lukas Fürst

**Observed rotations  
in just 100 μs**

**Real-time tracking of  
concentration with DCS!**



# NUV (350 nm) – CH<sub>2</sub>O



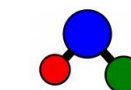
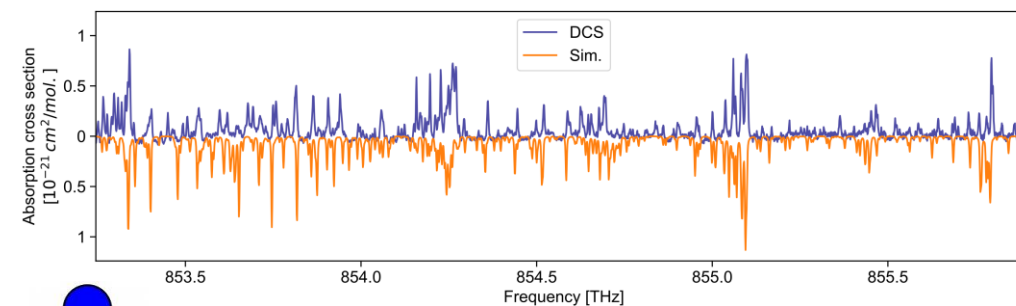
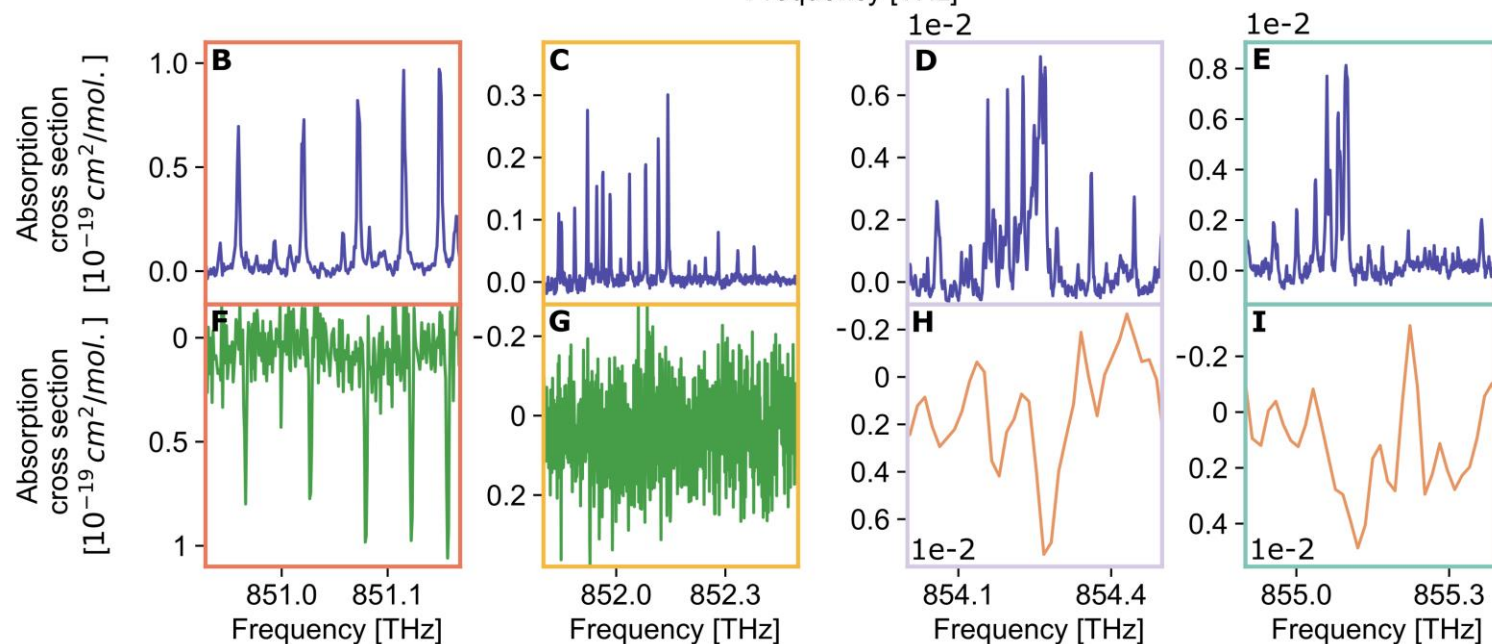
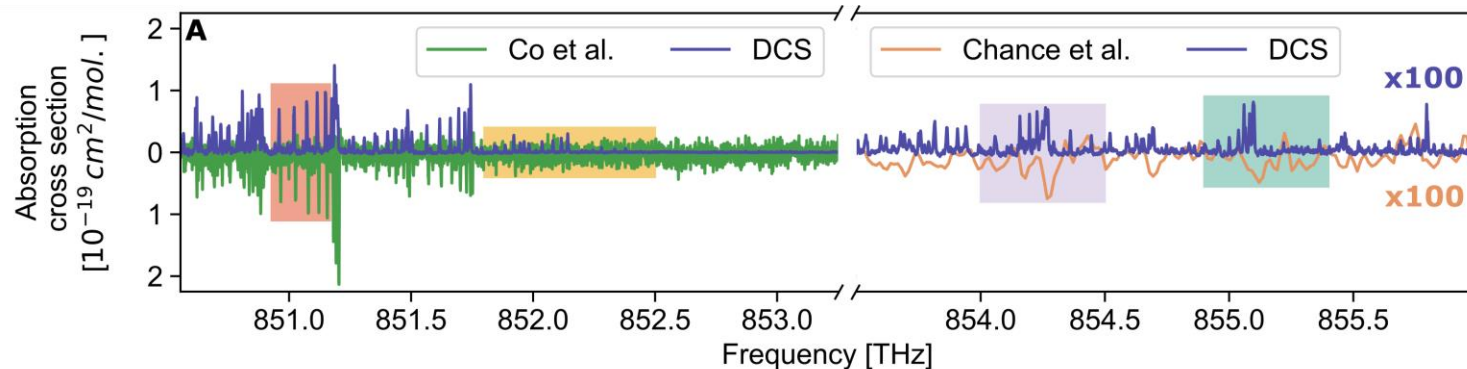
Lukas Fürst

1 GHz resolution in 500 ms  
(~ 0.03 cm<sup>-1</sup>)

Resolving power 10<sup>6</sup>

>100 prev. unresolved lines

Line assignments & simulation



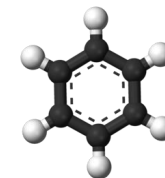
PGOPHER

DOI: 10.1016/j.jqsrt.2016.04.010

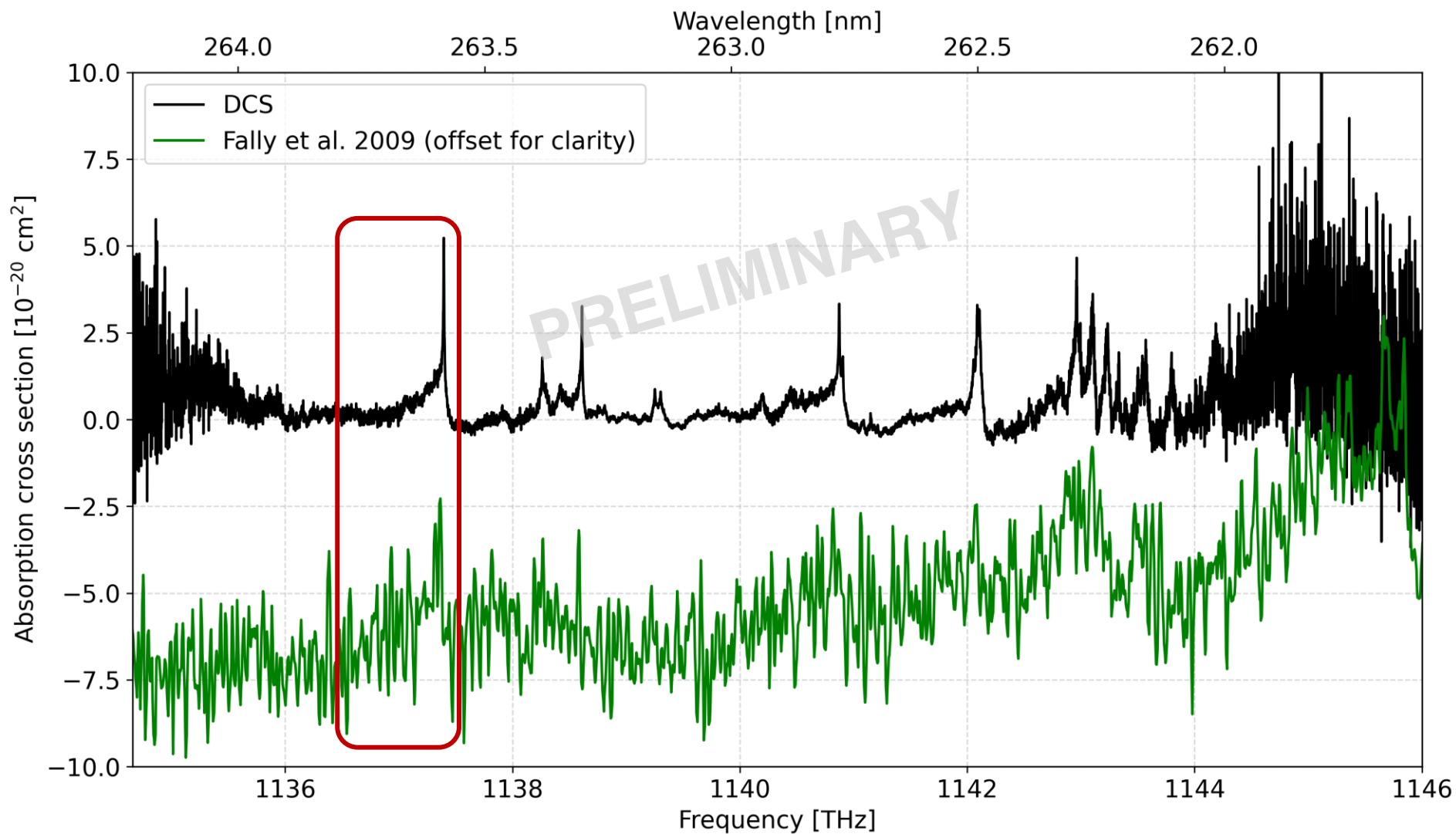
Lit: **Co et al. 2005**, DOI: 10.1021/jp053466i

Lit: **Chance et al. 2011**, DOI: 10.1016/j.jqsrt.2011.02.002

**Fuerst et al. submitted**, DOI: 10.48550/arXiv.2501.07350



# DUV (263 nm) – Benzene

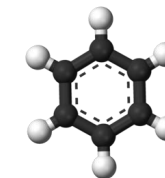


**Doppler-limited!**

**25x higher resolution**

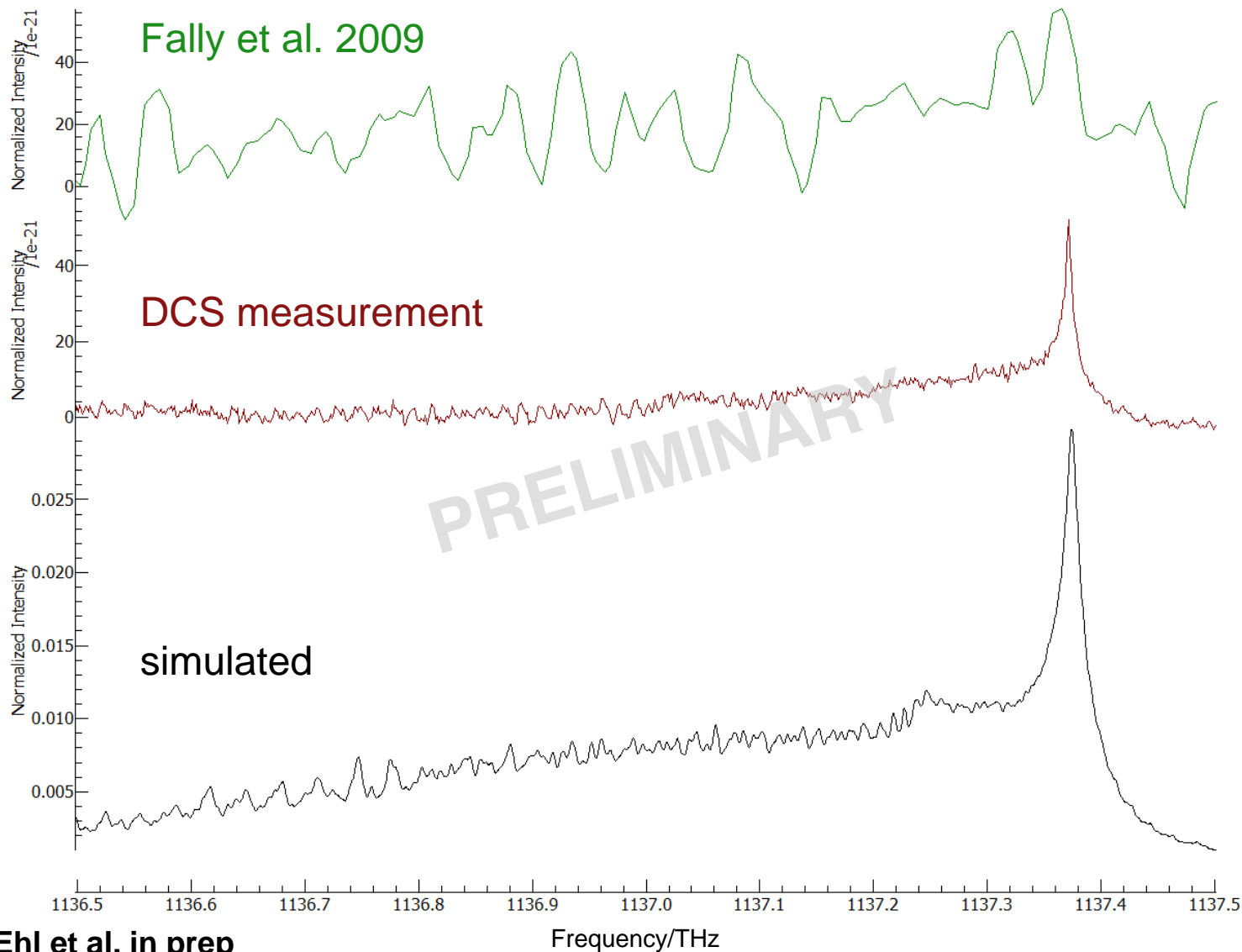
**$< 10^{-20} \text{ cm}^2$  sensitivity**

Doppler Width:  
1.5 GHz



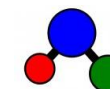
# DUV (263 nm) – Benzene

Fally et al. 2009



Several unassigned lines

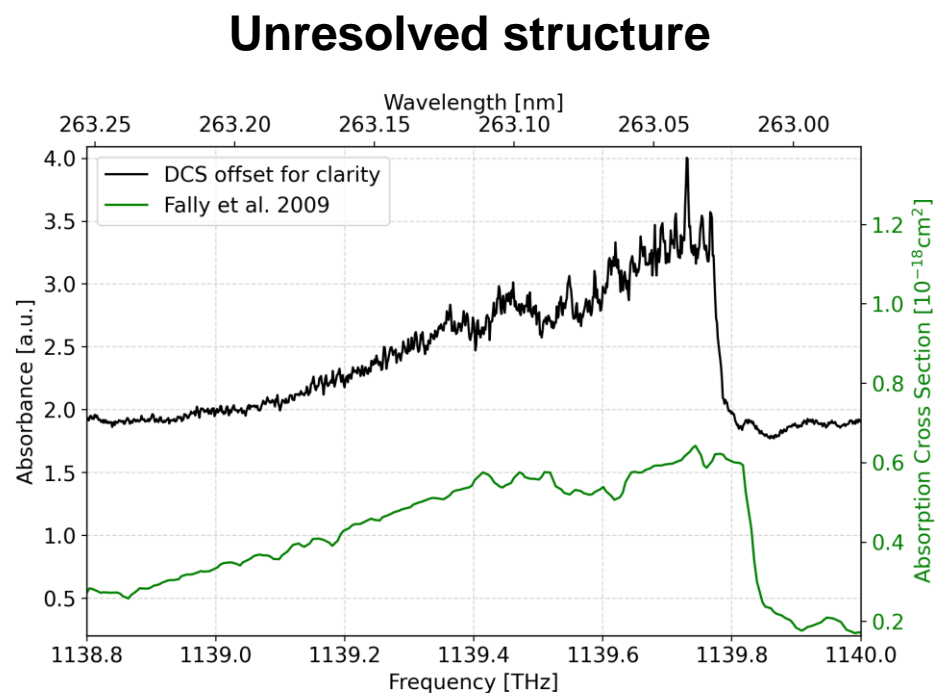
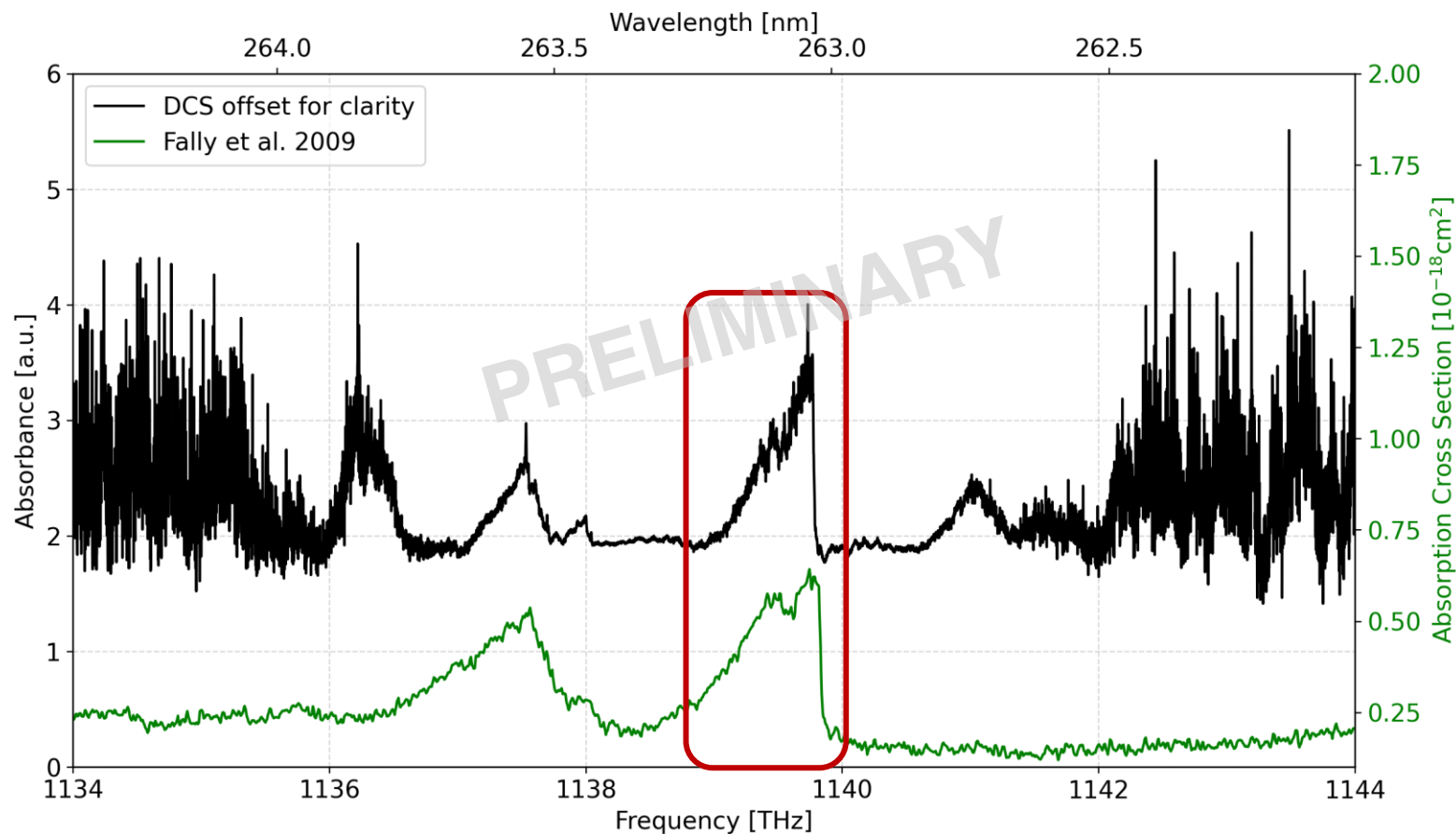
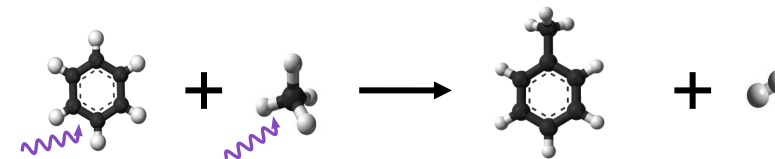
Ab initio calculations,  
line assignments and  
modelling in progress



PGOPHER

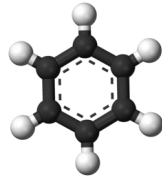
DOI: 10.1016/j.jqsrt.2016.04.010

# DUV (263 nm) – Toluene



# Outlook

- **Benzene details!**

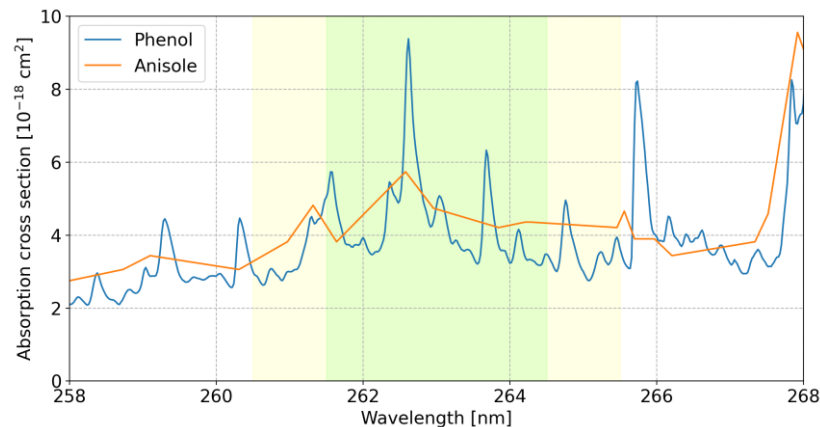


- Benzene-d6
- T dependence

- **New samples**

- SO<sub>2</sub>
- Aromatics

**SUGGESTIONS?**



- **HWO & Pollux**



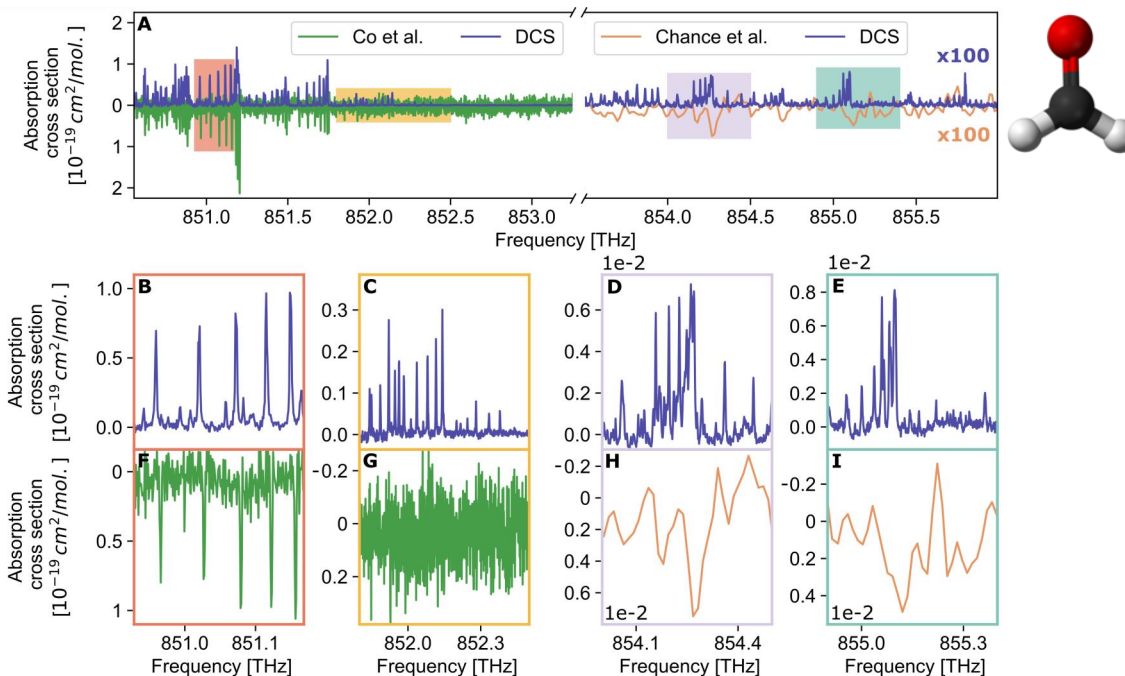
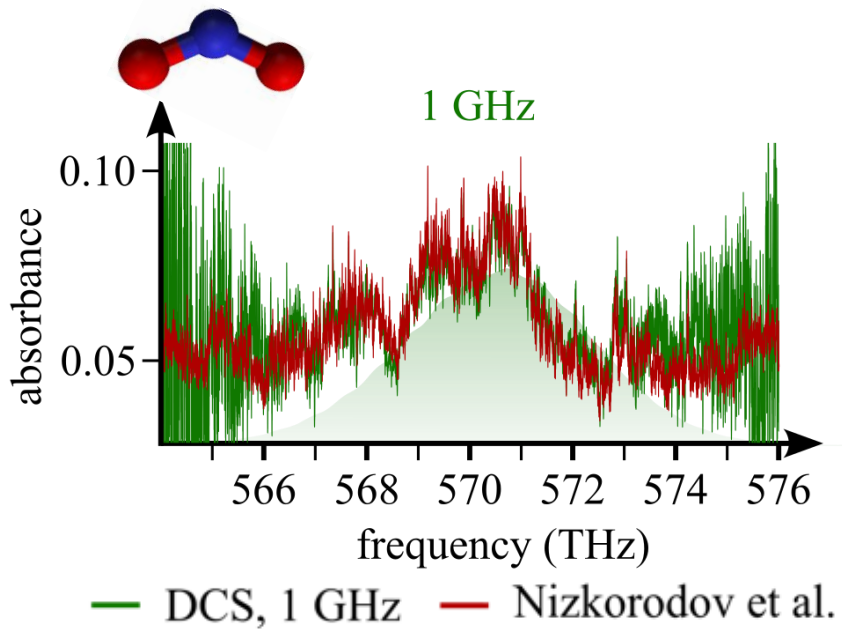
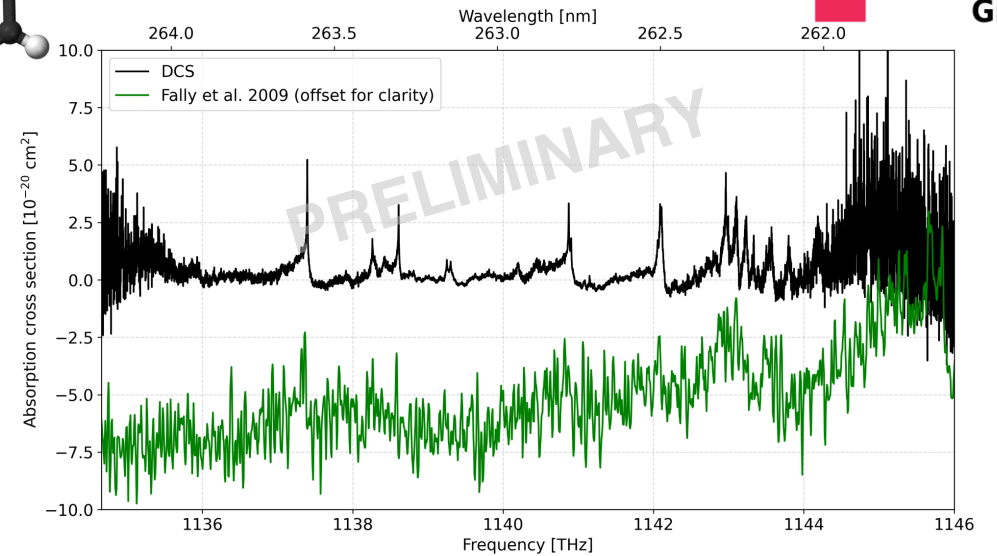
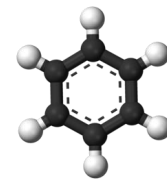
- Launch 2040
- Spectropolarimeter
- 100 – 1900 nm
- R ~ 100,000

- **Wishlist**

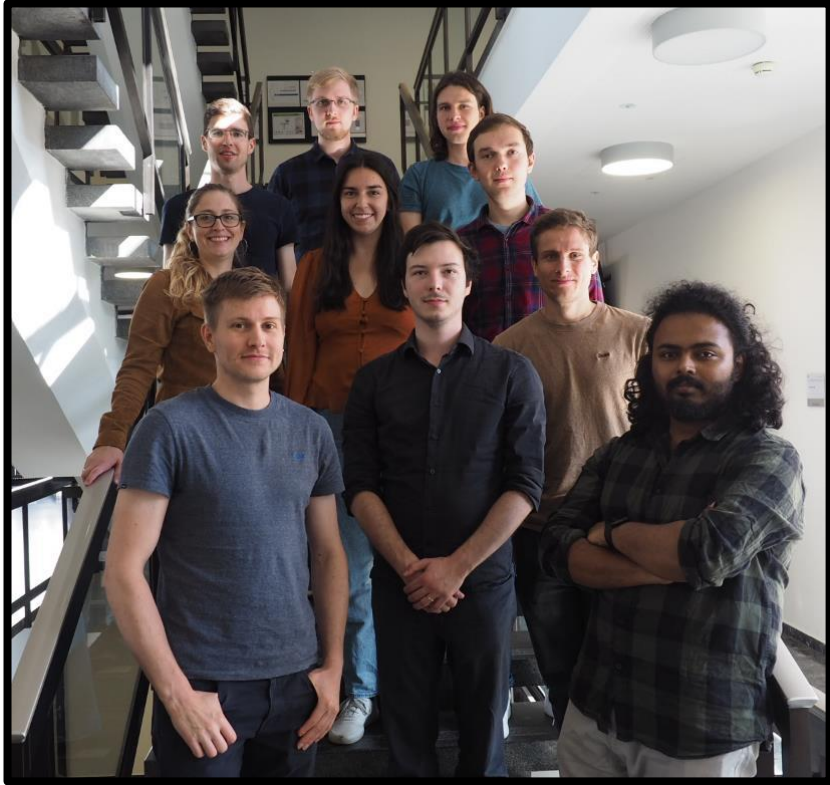
- SO<sub>2</sub> line list until 40,000 cm<sup>-1</sup>
- NO line list until 85,000 cm<sup>-1</sup>
- NO<sub>2</sub> & H<sub>2</sub>CO
- Benzene?

# Summary

- High spectral resolution ( $\sim$ MHz – GHz)
- Resolving power  $> 10^6$
- Fast acquisition times ( $\sim$ ms – s)



# Acknowledgements



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Florian Lindorfer

*Space Research Institute (IWF)*

**Prof. Dr. Luca Fossati**

*Funding*



# Extra Slides

# Setup Sketch

